

# **Entry, Descent, and Landing Systems Short Course**

Subject: Aerodynamics & Aerothermodynamics

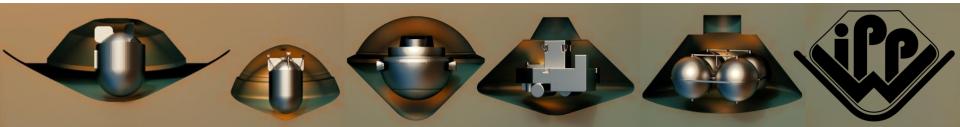
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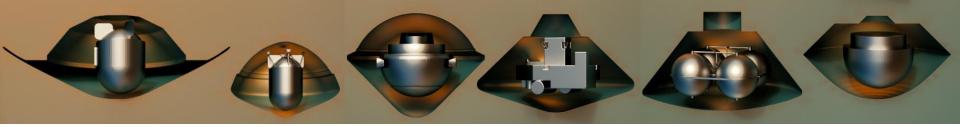
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International Planetary Probe Workshop 10
June 15-16, 2013
San Jose, California



## **Outline**

- Aerodynamics
  - Introduction/History/Methods
  - Dynamic Stability
- Aerothermodynamics
  - Introduction
  - Methods
  - Special Topics
- Case Study: MSL
- Q&A

Material focuses on blunt rigid capsules



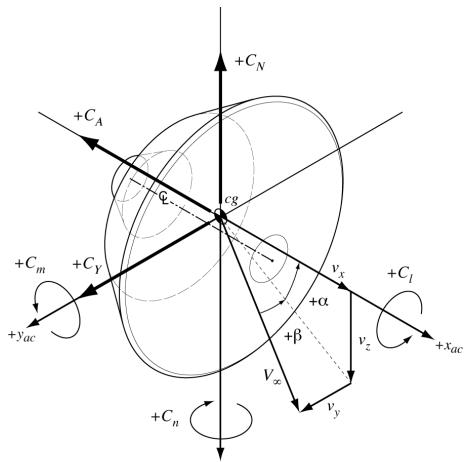
# Aerodynamics



### **Definitions**

Aerodynamics = forces & moments imparted on the entry vehicle by the atmosphere

Primary EDL Needs: High C<sub>D</sub>, Static stability ( $dC_m/d\alpha < 0$ )



Force Coefficient

$$C_X = \frac{F_X}{\frac{1}{2}\rho_{\infty}V_{\infty}^2 S_{ref}}$$

Moment Coefficient

$$C_X = rac{F_X}{rac{1}{2}
ho_\infty V_\infty^2 S_{ref}} \hspace{0.5cm} C_x = rac{M_x}{rac{1}{2}
ho_\infty V_\infty^2 S_{ref} d_{ref}}$$

$$F = \int_{surface} (pressure + shear)$$

$$M = \int_{surface} (pressure + shear) \times (l - l_{ref})$$

Lift Coefficient  $C_L = C_N \cos \alpha - C_A \sin \alpha$ 

Drag Coefficient  $C_D = C_N \sin \alpha + C_A \cos \alpha$ 

Pitch and Yaw Damping Coefficients

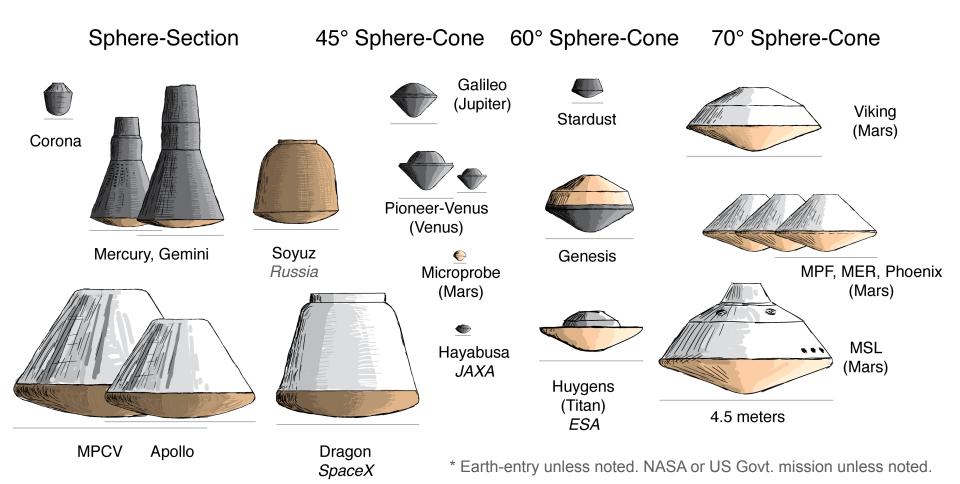
$$C_{m_q} + C_{m_{\dot{\alpha}}} = \frac{\partial C_m}{\partial \frac{qd_{ref}}{2V_{\infty}}} + \frac{\partial C_m}{\partial \frac{\dot{\alpha}d_{ref}}{2V_{\infty}}}$$

$$C_{n_r} - C_{n_{\dot{\beta}}} = \frac{\partial C_n}{\partial \frac{rd_{ref}}{2V_{\infty}}} - \frac{\partial C_n}{\partial \frac{\dot{\beta}d_{ref}}{2V_{\infty}}}$$

<sup>+</sup>zac International Planetary Probe Workshop 10 Short Course 2013

## The Blunt-Body Arsenal\* [2-37]

- EDL vehicles generally have blunt heatshields → high aerodynamic drag
- Other geometry drivers: packaging, stability, heating, science
- Backshells are generally selected to accommodate payload



# Flight Regimes (e.g. MSL)

#### Each flight regime requires unique prediction methods

ENTRY Parachute Deploy (V=5.8 km/s) (V=0.5 km/s)

Rarefied Hypersonic Supersonic Dynamics Transitional Supersonic 1000 > Kn > 0.00126.6 > Mach > 6 Knudsen > 10006 > Mach > 1.53.5 > Mach > 1.5LAURA (full aeroshell) w/ Ballistic Range Data LAURA CFD (forebody) DAC DSMC code DAC DSMC code base correction Molecules do Molecular collisions Chemically reacting flow Wake aerodynamics Unsteady wake introduces Peak heating, contribute to capsule capsule dynamic not collide modeled, still not Peak dynamic pressure continuum flow forces/moments instabilities Uncertainties for Lifting Vehicle RCS interactions ——— Higher L/D aerodynamics Aero-load deformation Balance Mass Jettison Lifting pitch/yaw Ablation shape change damping

Regime Definition

Data Source:

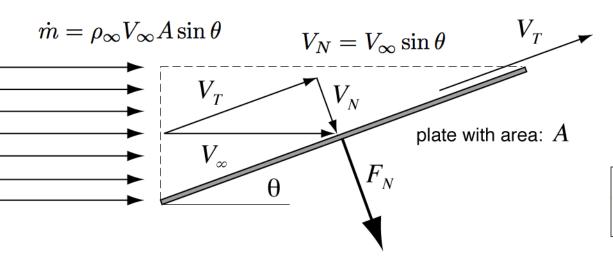
Characteristic Flow Physics

New Technical Challenges for MSL

# Newtonian Aerodynamics [38,39]

Newtonian method is an approximate method for estimating static aerodynamic coefficients

Newtonian method is more accurate for blunt bodies at hypersonic conditions



$$F = \dot{m}V_N = \rho_\infty A V_\infty^2 \sin^2 \theta$$

$$\frac{F}{A} = \rho_{\infty} V_{\infty}^2 \sin^2 \theta = p - p_{\infty}$$

$$C_p = rac{p - p_{\infty}}{rac{1}{2}
ho_{\infty}V_{\infty}^2} = rac{2V_N^2}{V_{\infty}^2} = 2\sin^2\theta$$

Lester Lees [40]: Updated Newtonian model replacing "2" with the maximum Cp based on Rayleigh Pitot equation (normal-shock relations), experiment or CFD.

Modified Newtonian "sine-squared law:"

$$C_p = C_{p_{max}} \sin^2 \theta$$

$$C_{p_{max}}=1.839$$
 for  $\gamma=1.4$ 

## Modified Newtonian Aerodynamics

#### From Lees [40]

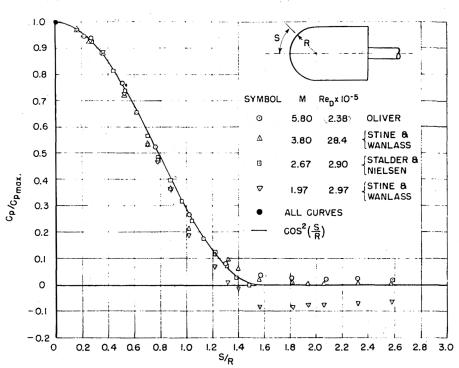


FIG.13- PRESSURE COEFFICIENT DISTRIBUTION ON SURFACE OF HEMISPHERE-CYLINDER AT VARIOUS MACH NUMBERS

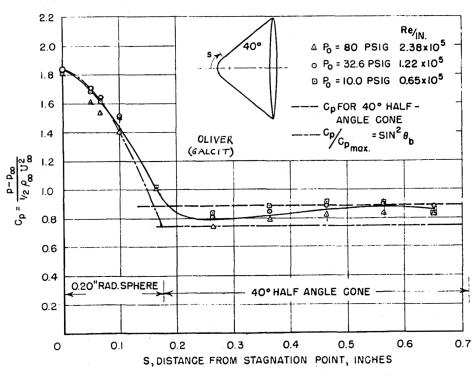
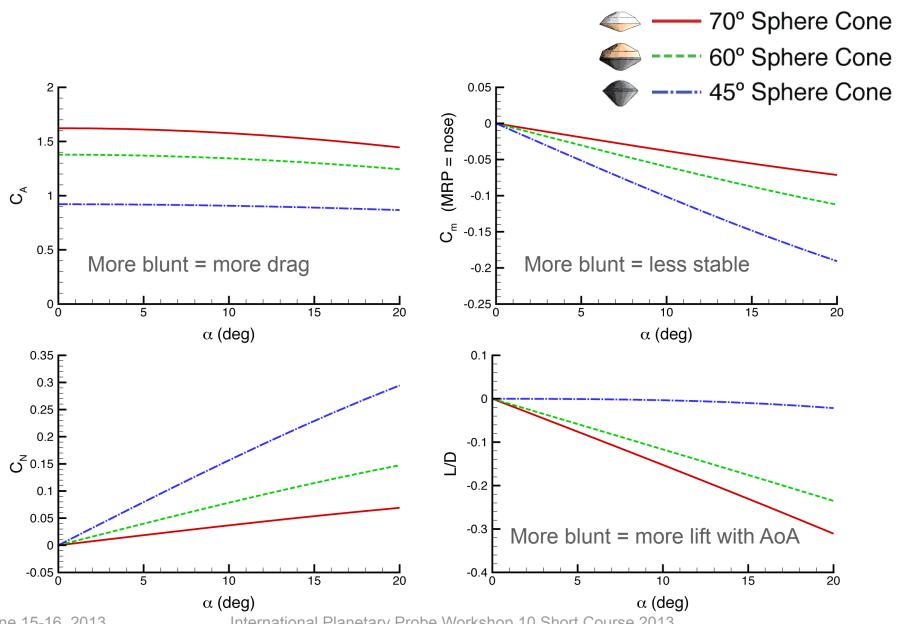


FIG.14 - PRESSURE DISTRIBUTION OVER BLUNT CONE ,  $M_{\infty}$  = 5.8 ,  $\alpha$  = 0 °

For hypersonic flows, local pressure is a function of the local surface angle relative to the flow

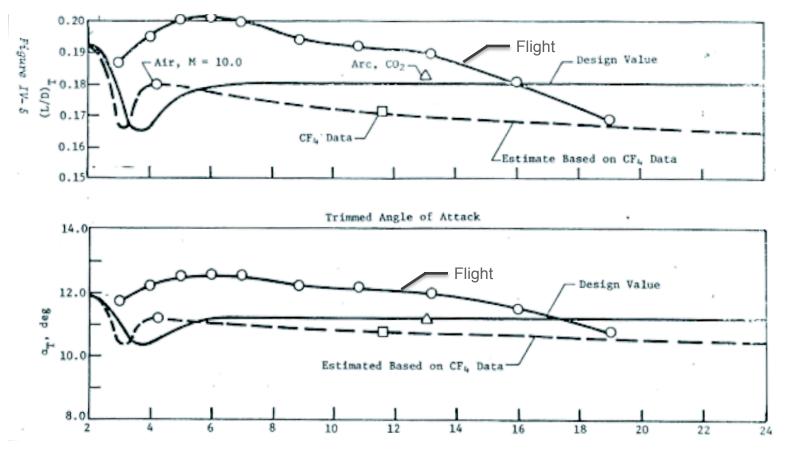
### Blunt Body Aerodynamic Characteristics

Modified Newtonian, Air, Mach = 20

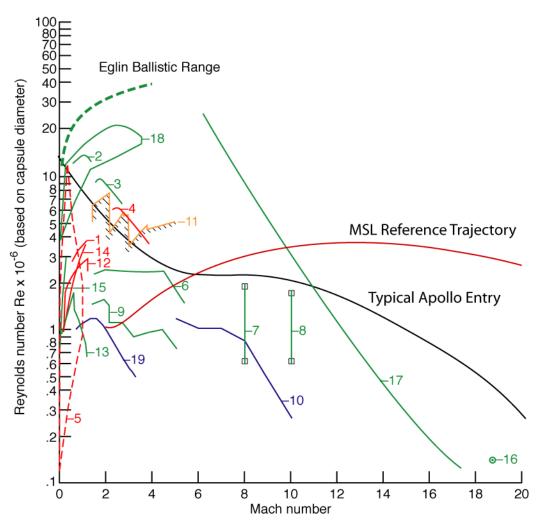


#### Viking Design vs. Flight Trim AOA and L/D [17]

- Viking was designed for L/D = 0.18 at a hypersonic trim AOA of 11 deg
- Aerodynamics were predicted from wind tunnel and ballistic range tests in air, CO2, CF4
- VL1 and VL2 trim AOA (and L/D) were higher than the design values
  - Believed at the time to be caused by an off-nominal CG location and outgassing/ ablation at the time, although low heat flux was experienced



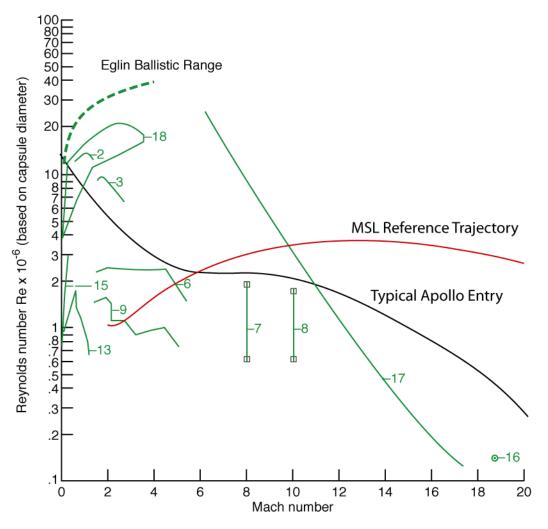
## National Aerodynamic Experimental Facilities



- 1. Ames 14-Ft Transonic wind Tunnel (0.055 scale) Ames Unitary Plan Wind Tunnels (0.105 scale)
- 2. 11 by 11 ft
- 3. 9 by 7 ft
- 4. 8 by 7 ft
- 5. Ames 12-Ft Pressure Tunnel (0.10 scale)
- 6. Arnold Engineering Development Center, von Karman Facility 50-In. Tunnel A (0.045 scale)
- 7. Arnold Engineering Development Center, von Karman Facility 50-In. Tunnel B (0.045 scale)
- 8. Arnold Engineering Development Center, von Karman Facility 50-In. Tunnel C (0.045 scale)
- 9. Jet Propulsion Laboratory 20-In. Supersonic Wind Tunnel (0.045 scale)
- 10. Jet Propulsion Laboratory 21-In. Hypersonic Wind Tunnel (0.02 scale)
- 11. Langley Unitary Plan Wind Tunnel (0.055 scale)
- 12. Langley 8-Ft Transonic Pressure Tunnel (0.055 scale)
- 13. Langley 16-Ft Transonic Dynamics Tunnel (0.08 scale)
- 14. Langley 16-Ft Transonic Wind Tunnel (0.085 scale)
- North American Aerodynamics Laboratory
   by 10-ft Low-Speed Wind Tunnel (0.105 scale)
- Arnold Engineering Development Center, von Karman Facility 50-In. Hot-Shot II, Tunnel H (0.04 scale)
- 17. Cornell Aeronautical Laboratory 48-in. Shock Tunnel (0.05 scale)
- 18. North American Aviation 7- by 7-ft Trisonic Wind Tunnel (0.105 scale)
- 19. North American Aviation Supersonic Aerophysics Laboratory (0.02 scale)

From [10] NASA TN-D 3748, "Apollo Wind Tunnel Testing Program Historical Development of General Configurations"

## National Aerodynamic Experimental Facilities



Ames Unitary Plan Wind Tunnels (0.105 scale)

- 2. 11 by 11 ft
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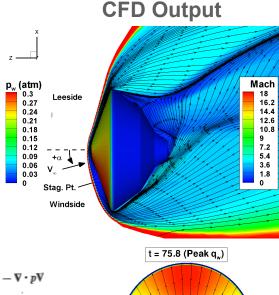
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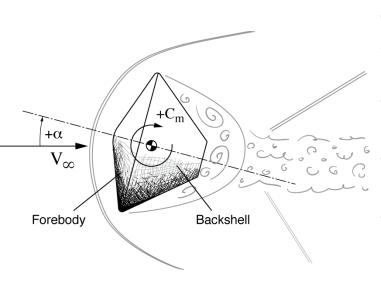
# Computational Fluid Dynamics (CFD)

- Definition: Numerical solution of the fluid dynamic equations of motion
- The closing of wind tunnels & advancement of computers has pushed CFD to the forefront of aerothermodynamics prediction
  - CFD is a high-fidelity method of "simulating" flight conditions (esp. hypersonic)
- But, the resources required can be large (people, time, computers), especially for complex 3-D geometries
- Examples: LAURA (NASA Langley), DPLR (NASA Ames)

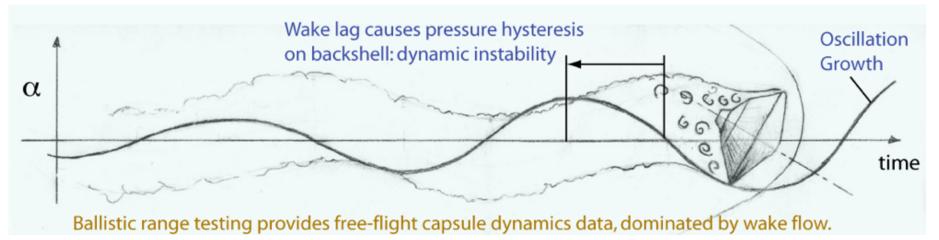
#### **Navier-Stokes Equations Vehicle Geometry CFD Grid** p<sub>w</sub> (atm) 0.3 Leeside 0.27 $\frac{\partial \rho}{\partial t} + \nabla \cdot (\rho \mathbf{V}) = 0$ 0.24 Continuity equation $\rho \frac{Du}{Dt} = -\frac{\partial p}{\partial x} + \frac{\partial \tau_{xx}}{\partial x} + \frac{\partial \tau_{yx}}{\partial y} + \frac{\partial \tau_{zx}}{\partial z}$ x Momentum $\rho \frac{Dv}{Dt} = -\frac{\partial p}{\partial v} + \frac{\partial \tau_{xy}}{\partial x} + \frac{\partial \tau_{yy}}{\partial v} + \frac{\partial \tau_{zy}}{\partial z}$ y Momentum Windside $\rho \frac{Dw}{Dt} = -\frac{\partial p}{\partial z} + \frac{\partial \tau_{xz}}{\partial x} + \frac{\partial \tau_{yz}}{\partial y} + \frac{\partial \tau_{zz}}{\partial z}$ z Momentum Energy $\rho \frac{D(e + V^2/2)}{Dt} = \rho \dot{q} + \frac{\partial}{\partial x} \left( k \frac{\partial T}{\partial x} \right) + \frac{\partial}{\partial y} \left( k \frac{\partial T}{\partial y} \right) + \frac{\partial}{\partial z} \left( k \frac{\partial T}{\partial z} \right) - \nabla \cdot p \nabla$ $+\frac{\partial(u\tau_{xx})}{\partial x}+\frac{\partial(u\tau_{yx})}{\partial y}+\frac{\partial(u\tau_{zx})}{\partial z}+\frac{\partial(v\tau_{xy})}{\partial x}+\frac{\partial(v\tau_{yy})}{\partial y}$ $+\frac{\partial(v\tau_{xy})}{\partial x}+\frac{\partial(w\tau_{xz})}{\partial x}+\frac{\partial(w\tau_{yz})}{\partial y}+\frac{\partial(w\tau_{zz})}{\partial z}$



## **Dynamic Stability**

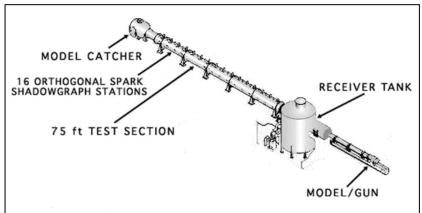


- Blunt bodies are dynamically unstable at supersonic speeds - driven by wake flow
- Phenomenon starts around Mach 3.5 and slower
- Stability can vary dramatically with backshell shape
- Unsteady CFD has provided insight on flow mechanisms
- Experiment is the only validated method of determining pitch damping coefficients



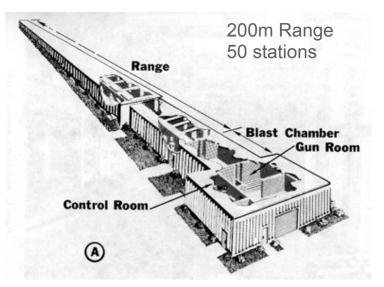
## Free Flight Dynamic Stability Test Facilities

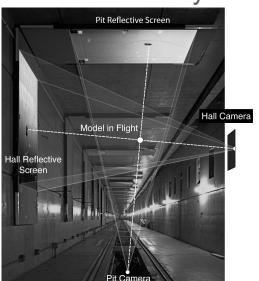
NASA Ames HFFAF (can test lifting, but Re is typically low)

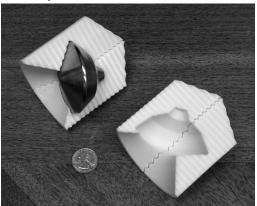




Eglin Air Force Base: Aeroballistic Test and Evaluation Facility (ATEF) (Used for all recent robotic missions but currently mothballed)

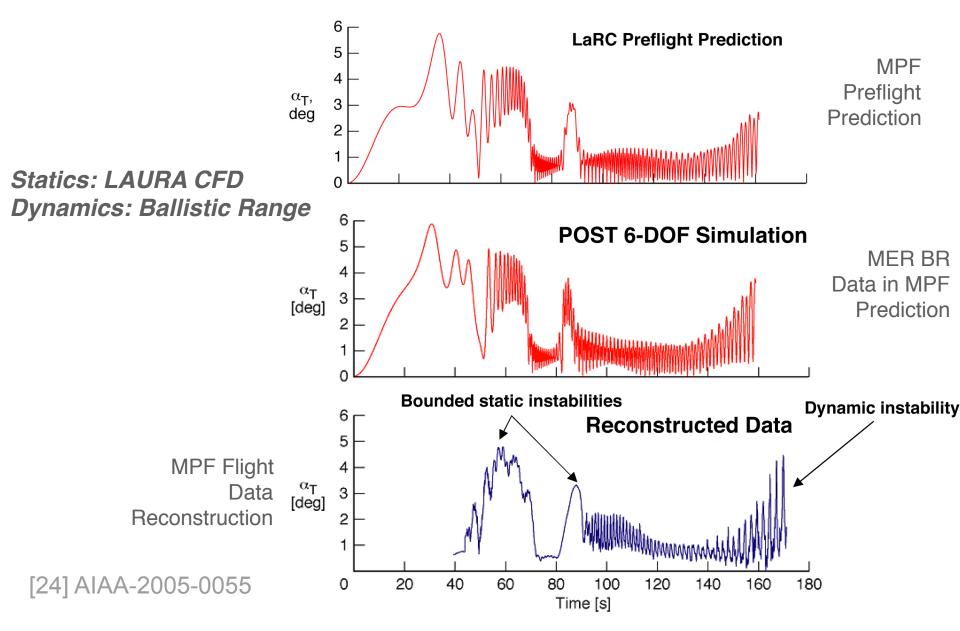


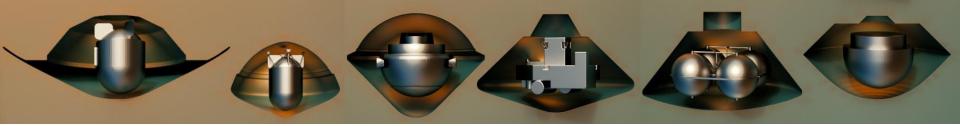




MSL model

#### Mars Pathfinder Reconstruction





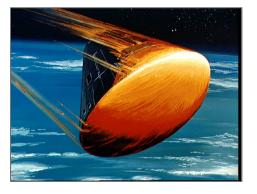
# Aerothermodynamics



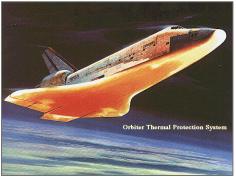
## What is Aerothermodynamics?

- Definition: Aerodynamic heating of a solid surface by a gas through a viscous boundary layer & high-temperature shock layer
  - First addressed in the 1950s with the advent of hypersonic re-entry missiles
- Aerodynamic heating is most often associated with hypersonic atmospheric flight → high velocity + dense atmospheric gas = high temperatures





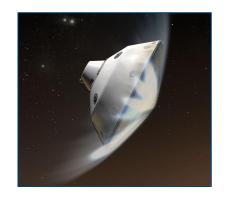
**Space Shuttle** 



Pioneer Venus



**Mars Science Laboratory** 

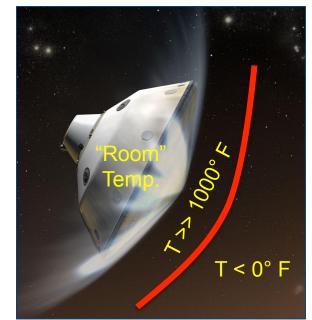


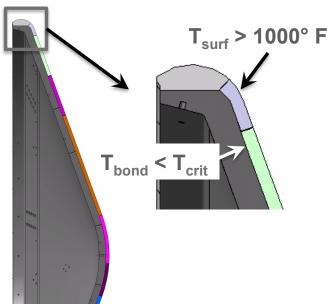
"Two major problems encountered today in aeronautics are the determination of skin friction and skin temperatures of high-speed aircraft."

E. R. Van Driest, 1950

## **Thermal Protection**

- Entry vehicle kinetic energy → increased shock layer temperature → increased surface temperature
- Payloads, human or robotic, must be protected from extreme temperatures
- Load-bearing structures cannot perform adequately at elevated temperatures
  - Material performance degrades
- The type & amount of thermal protection system (TPS) material depend on the prediction of aerodynamic heating at flight conditions
  - Heating = f(time), surface pressure & shear stress may also be important
  - Difficult/impossible to simulate flight conditions in ground facilities
- TPS adds mass to the EDL system!



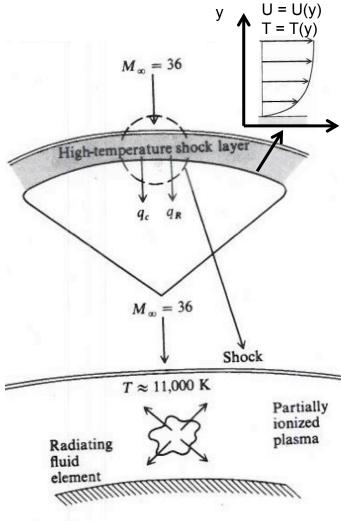


## **Terminology**

- <u>Convection</u> = heat transfer via <u>conduction</u> & <u>diffusion</u> through a <u>viscous boundary layer</u>
  - Conduction ~ Temperature gradient at surface
  - Diffusion (aka Catalytic Heating) ~ Gas/surface chemical reactions
- Radiation = heat transfer via atomic excitation in a high-temperature shock layer
  - Radiation ~ Shock layer temperature

# Convection Radiation $q_{w} = q_{Cond} + q_{Diff} + q_{Rad}$

- Heat rate  $(q_w)$  = instantaneous heat transfer  $(W/cm^2)$
- Heat load  $(Q_w = \int q_w dt)$  = integration of  $q_w$  (J/cm<sup>2</sup>)
- Convective & radiative heating magnitudes depend on the entry vehicle & conditions
  - Mars: mostly convective
  - Earth (lunar return), Venus, Jupiter: convective & radiative



Schematic of Aerodynamic Heating to a Blunt Body (Ref. Anderson)

# **Engineering Methods**

- Definition: Analytical solutions to the boundary layer equations
- Entry vehicles can be approximated by simple geometries
  - Nose (hemisphere), leading edge (cylinder), surface (cylinder or flat plate)
- Engineering methods have an important place in conceptual vehicle design
  - Quick to use, based on theory, can be integrated into trajectory codes
- Examples for stagnation point heating:
  - Requires effective nose radius (R<sub>n</sub>)

Fay & Riddell (1958) 
$$\dot{q}_w = \frac{0.763}{\left(\text{Pr}_w\right)^{0.6}} \left(\rho_e \mu_e\right)^{0.4} \left(\rho_w \mu_w\right)^{0.1} \left[\left(h_o\right)_e - h_w\right] \left[1 + \left(Le^{0.52} - 1\right) \frac{h_d}{\left(h_o\right)_e}\right] \left[\frac{du_e}{dx}\right]_t$$

Sutton & Graves (Convective) 
$$q_c = C \left(\frac{\rho_{\infty}}{R_n}\right)^{\frac{1}{2}} V_{\infty}^3$$
 Earth: C = 1.7415e-4 Mars: C = 1.9027e-4

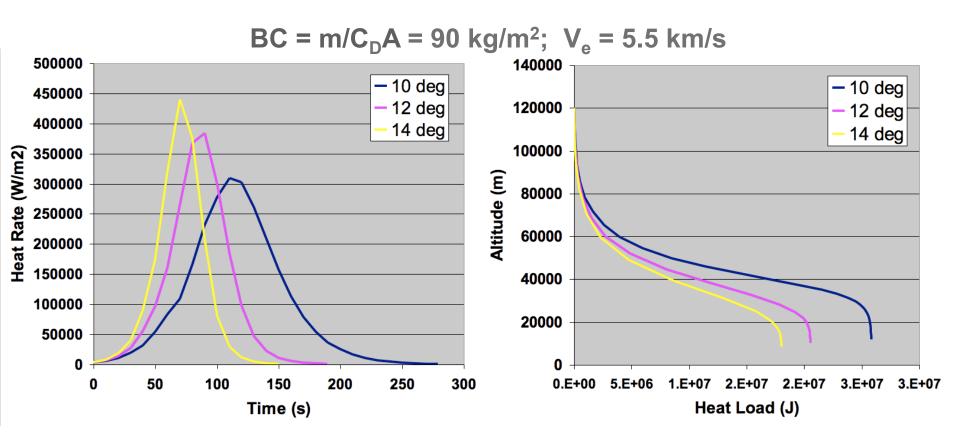
**Simplified Entry Vehicle Geometries** 

$$q_r = C_i R_n^a \rho_{\infty}^m f_i(V_{\infty})$$

 $q_r = C_i R_n^a \rho_\infty^m f_i(V_\infty)$  Earth: a = 1, m = 1.2 Mars: a = 0.526, m = 1.2  $\sim 1/R_n$ 

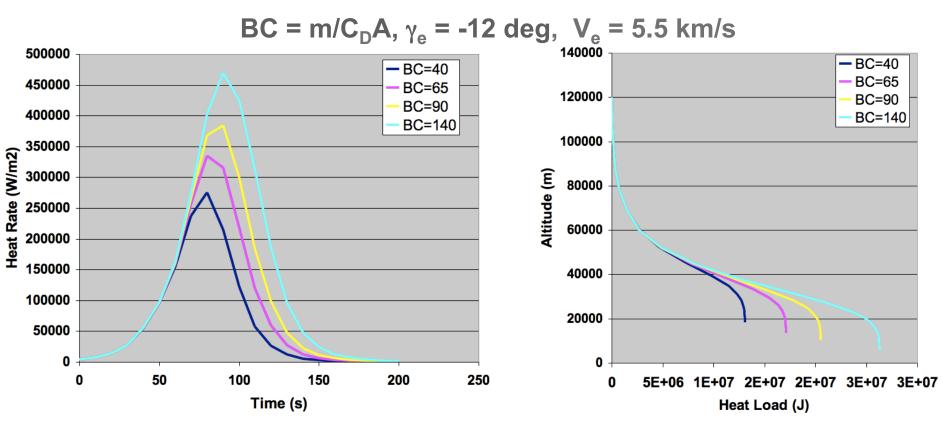
## Example: Mars, Effect of Flight Path Angle

- Assumptions: Sutton-Graves stagnation point heating ( $R_n = R_{eff}$ ), ballistic entry, exponential atmospheric density
- A steeper FPA <u>increases</u> max. heat flux, but <u>decreases</u> heat load (TPS thickness)
  - Potential for TPS mass savings by using a steeper entry



### Example: Mars, Effect of Ballistic Coefficient

- Assumptions: Sutton-Graves stagnation point heating ( $R_n = R_{eff}$ ), ballistic entry, exponential atmospheric density
- If the BC increases, heating will increase
  - Example: payload mass grows, but aeroshell diameter does not

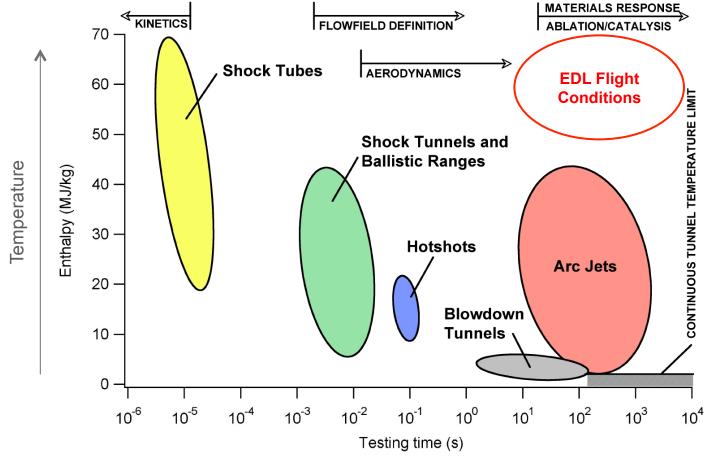


## Rules of Thumb for Ballistic Entry

- Stagnation point convective heating:  $q_c \sim \rho_{\infty}^{1/2} V_{\infty}^{3} / R_n^{1/2}$
- Stagnation point radiative heating: q<sub>r</sub> ~ R<sub>n</sub><sup>m</sup>
- Wall temperature: T<sub>w</sub><sup>4</sup> ~ q<sub>c</sub>
- Convective heat flux, q<sub>c</sub>:
  - ↑ with ↑ entry velocity
  - → with → ballistic coefficient (mass)
  - ↑ with ↑ entry flight path angle
- Convective heat load:  $Q_c = \int q_c dt$ 
  - ↑ with ↑ entry velocity
  - → with → ballistic coefficient (mass)
  - → with ♥ entry flight path angle

# Ground-to-Flight Traceability

- Ground-based testing, while valuable, cannot replicate all EDL flight conditions
  - Technically & fiscally prohibitive to build & run such facilities
- Aerodynamic heating analysis at flight conditions is usually left to engineering methods (conceptual) & high-fidelity CFD models (TPS design)
  - Ground data are used to anchor models to be used for flight predictions



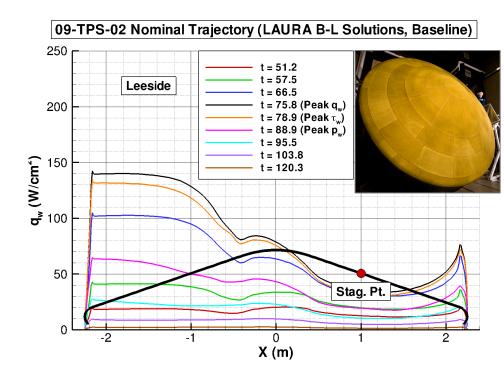
## Ground-to-Flight Traceability

- Heating magnitudes in ground facilities ≠ heating in flight
  - Differences in Mach, Re, gas composition…
- Ground facilities are used to understand qualitative heating and to provide a source of data for CFD validation
  - Ground testing does not generally include TPS response, which is added post-CFD

#### MSL Testing in AEDC Tunnel 9 (Mach 8)

## AEDC T9 Run 3048 (Mach = 8, Re<sub>...n</sub> = 24.8 x $10^6$ , $\alpha$ = 16 deg) 1.4 .AURA B-L 1.2 Data +/- 12% 0.4 0.2 X (in)

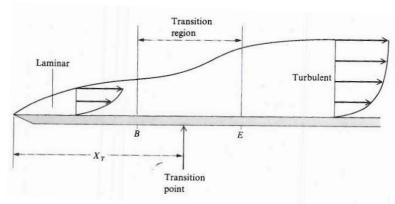
#### MSL Flight CFD (30 > Mach > 7)



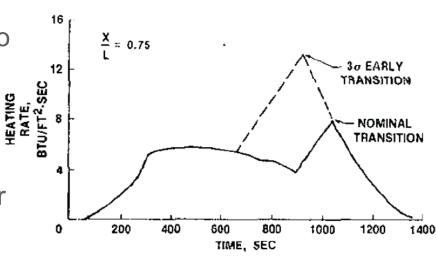
# **Boundary Layer Transition (BLT)**

- Predicting BLT is one of the most difficult aspects of aerodynamic heating
  - Mechanisms not well understood
- BLT can increase max. heat flux and total heat load → TPS mass & performance
  - $-q_{Turb} > or >> q_{Lam}$
  - Stagnation point may not have highest q<sub>w</sub>
- Engineering methods are often used to predict BLT timing
  - BL momentum-thickness Reynolds no.,  $Re_{\theta} > Re_{\theta,crit}$
  - Examples: Space Shuttle, MSL
- Conservative approach is to design for turbulent conditions using CFD

"When I meet God, I am going to ask him two questions: Why relativity? And why turbulence? I really believe he will have an answer for the first." - Werner Heisenberg



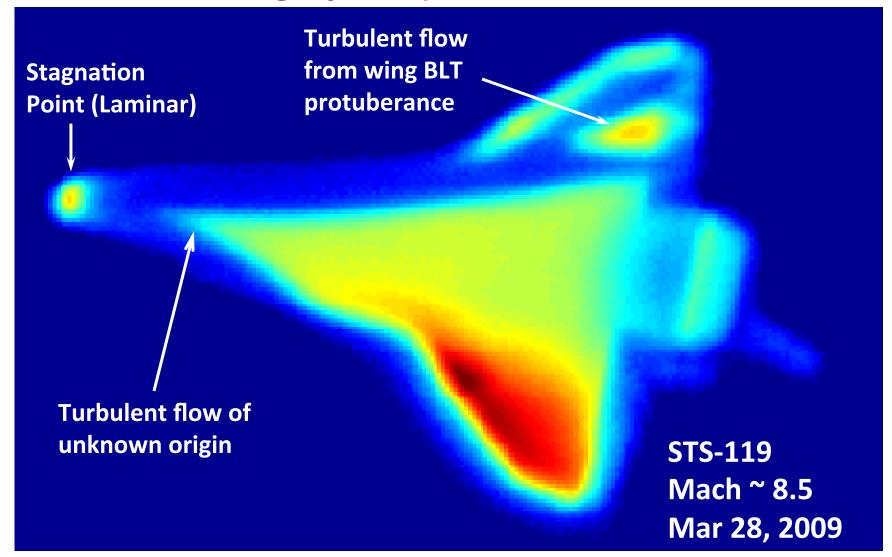
Schematic of BLT (Ref. Anderson)



BLT Design for Space Shuttle Windward Body Point (Ref. AIAA 79-1042)

#### BLT on STS-119

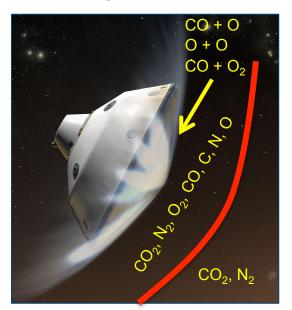
HYTHIRM imagery of Space Shuttle lower surface



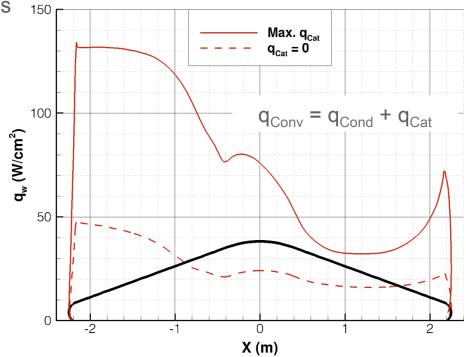
## Catalytic Heating

- High shock-layer temperature dissociates gas molecules
  - Mars:  $CO_2/N_2$  in front of shock  $\rightarrow CO_2$ ,  $N_2$ ,  $O_2$ , CO, C, N, O behind shock
  - Earth:  $N_2/O_2$  in front of shock  $\rightarrow N_2$ ,  $O_2$ , N, O, NO + ions behind shock
- Post-shock gas may recombine at surface
  - Heat of reaction = Catalytic heating = Higher convective heating
- Conservative approach is to maximize catalytic heating
  - "Super-catalytic" = recombine all atoms

#### **Mars Non-Equilibrium Chemistry**



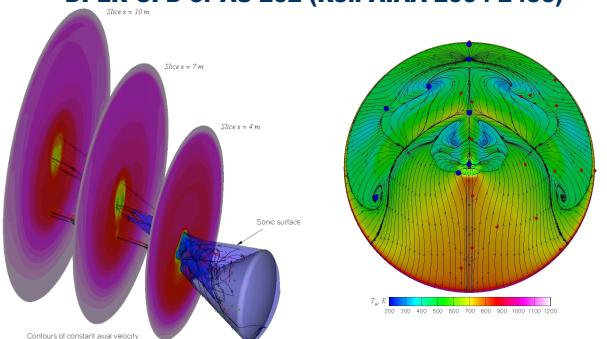
# MSL Predicted Convective Heat Flux at Peak Heating (No Uncertainties)



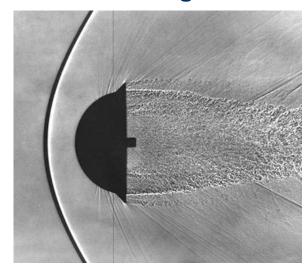
## Aftbody Heating

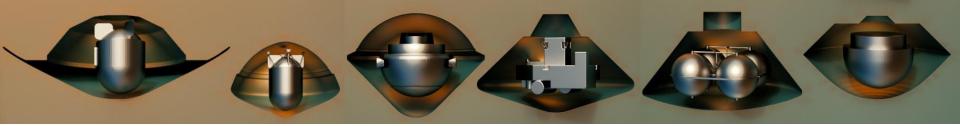
- Heating on vehicle components that are in the wake of a hypersonic flowfield are more difficult to predict
  - Separated, unsteady, turbulent, low-density flow
  - Uncertainties are > than for heatshield
- Heating is generally < 10% of stagnation point heating for a blunt body

#### **DPLR CFD of AS-202 (Ref. AIAA 2004-2456)**



#### **Ballistic Range Model**





# Case Study: Mars Science Laboratory

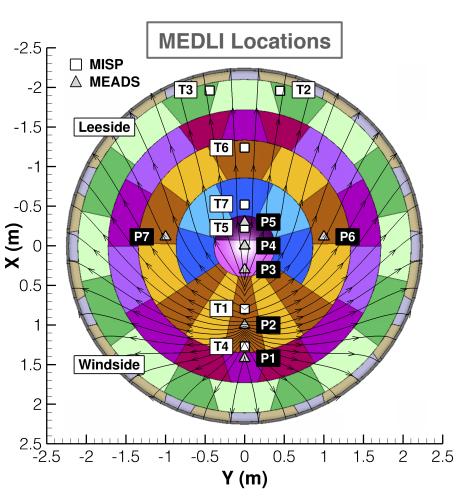
## Mars Science Laboratory (MSL)

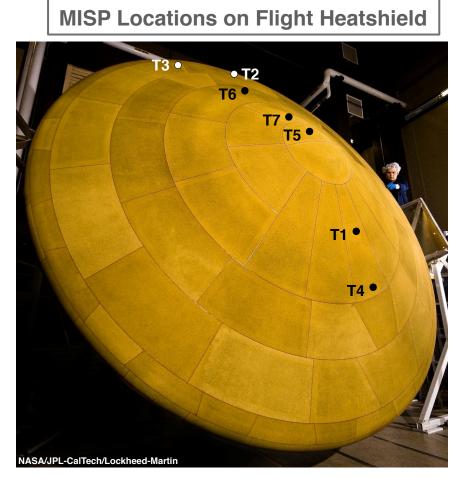
- Guided entry (trim  $\alpha$  = -16 deg, trim L/D = 0.24), RCS bank control
- Large heatshield + high m/C<sub>D</sub>A = BLT expected prior to peak heating, resulting in high heating rate

	Viking 1/2	Pathfinder	MER A/B	Phoenix	MSL
Diameter, m	3.5	2.65	2.65	2.65	4.5
Entry Mass, kg	930	585	840	602	3152
Entry Vel., km/s	4.5/4.42	7.6	5.5	5.9	5.9
Entry FPA, deg	-17.6	-13.8	-11.5	-13	-16.1
$m/(C_DA)$ , $kg/m^2$	64	62	90	65	135
Hypersonic $\alpha$ , deg	-11.2	0	0	0	-16
Hypersonic L/D	0.18	0	0	0	0.24
Heat Flux, W/cm <sup>2</sup>	24	106	48	56	>200 (Design)

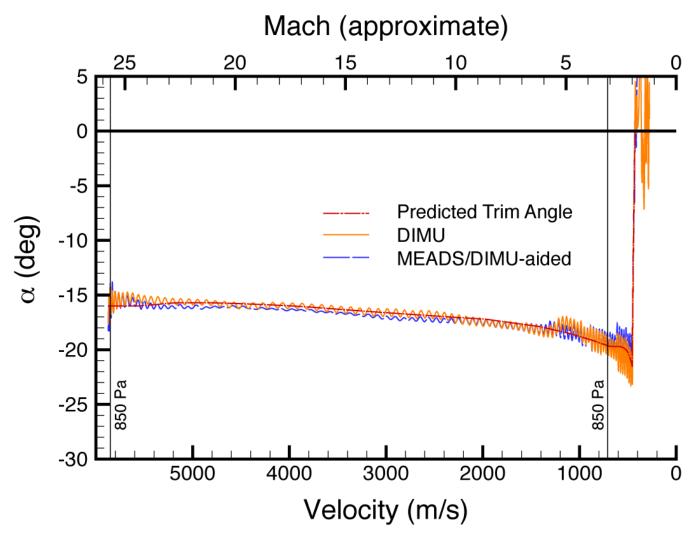
#### MSL Heatshield & MEDLI Instrumentation

- MISP: 2-4 thermocouples (0.1-0.7 in below surface) at 7 locations
  - Surface heating comes from inverse analysis of temperature data
- MEADS: pressure transducer at 7 separate locations → capsule attitude



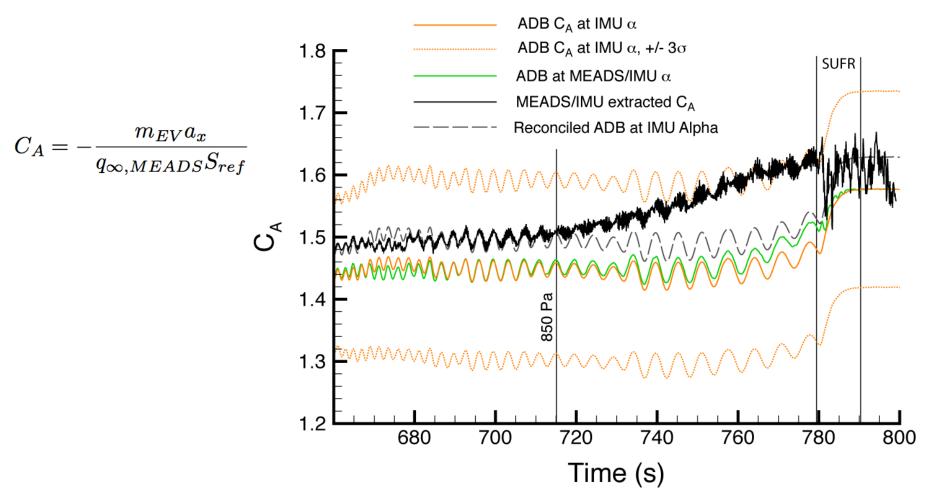


## MSL Reconstructed Angle-of-Attack [28]



Trim angle and Lift-to-Drag predictions were very accurate

# $MSLC_A[28]$

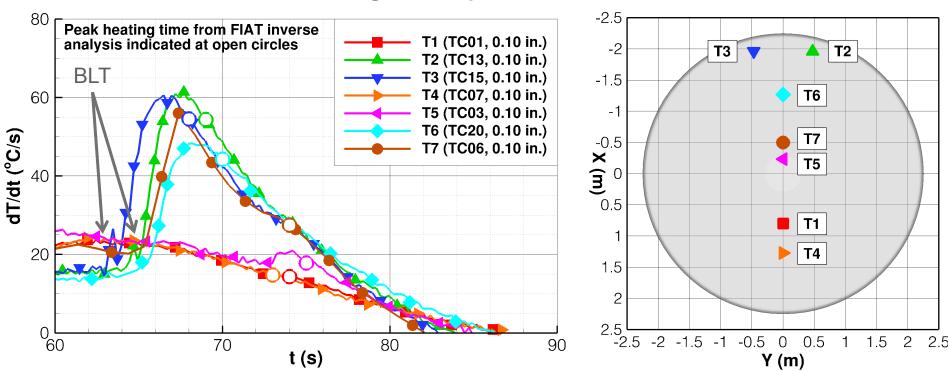


- MEADS Supersonic C<sub>A</sub> remains area of study
- •Viking-derived base pressure correction remains possible error source
- Winds and supersonic drag are the two dominant contributors to landing miss distance

## **MSL Boundary Layer Transition**

- BLT indicated by rapid increase in dT/dt
- MISP T2, T3, T6, and T7 all experienced BLT
- Small bump at T5 believed to indicate BLT
- All BLT events occurred prior to peak heating, as expected

#### Time Rate-of-Change of Temperature 0.1 in. Below Surface



# MSL Reconstructed Heating

### Expected results:

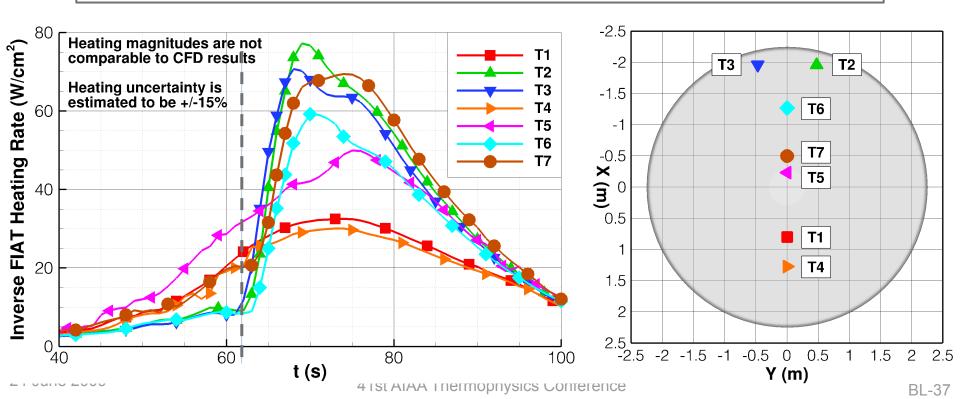
Laminar heating highest at T5
Laminar heating lowest at T2, T3, T6
Turbulent heating highest at T2, T3

### Unexpected results:

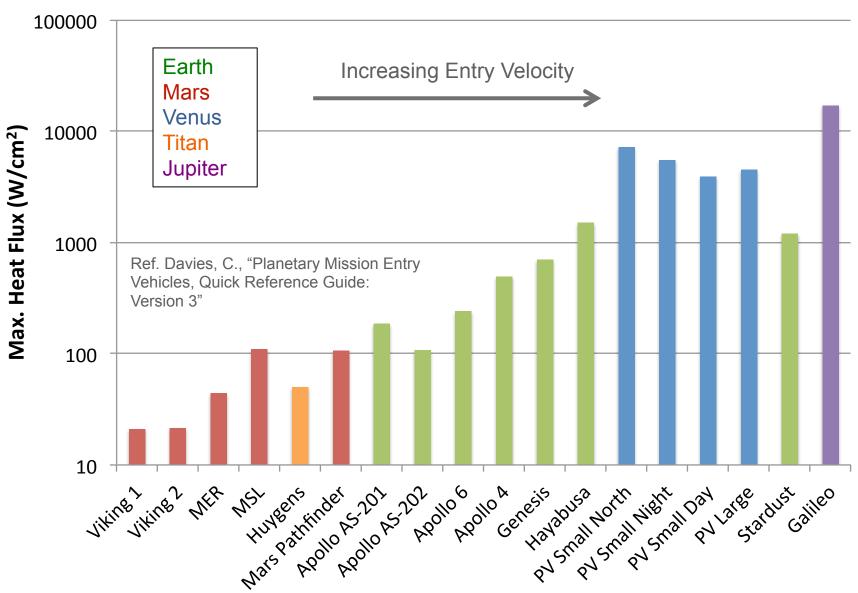
High turbulent heating at T7 relative to T2, T3  $\rightarrow$  roughness?

High heating at T1, T4 relative to T2,  $T3 \rightarrow radiation$ ?

#### MISP Heating Rate from FIAT Inverse Analysis (Estimated +/-15% Uncertainty)

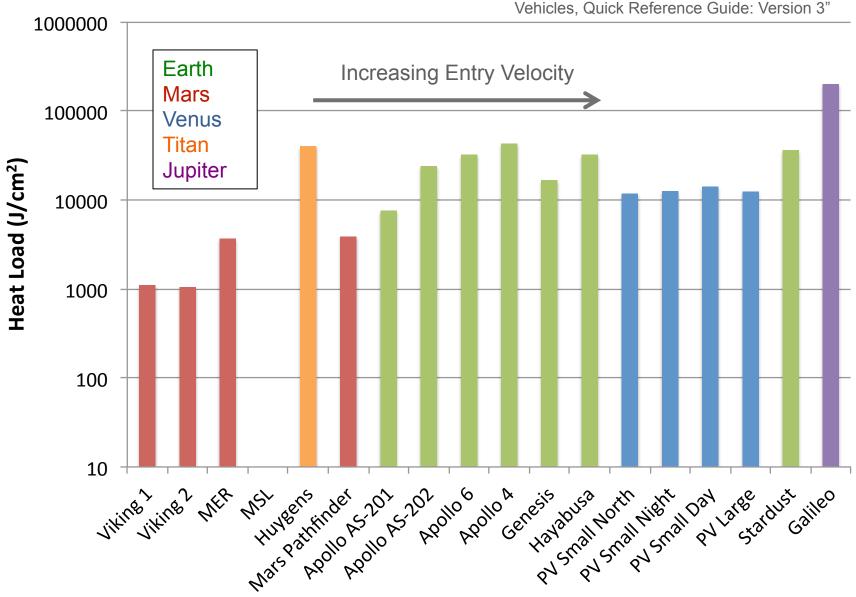


# Max. Heat Flux for Select Entry Vehicles

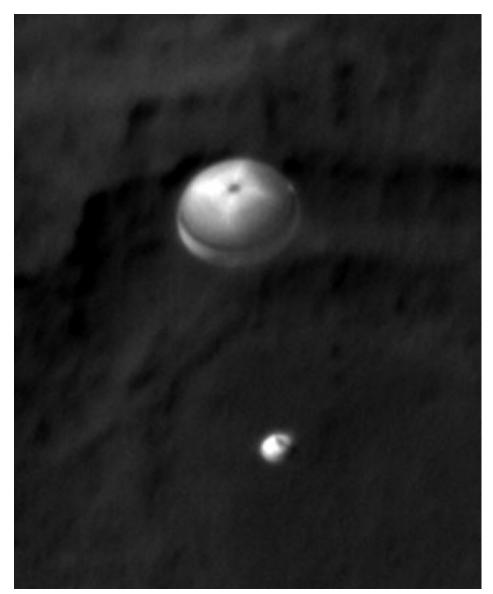


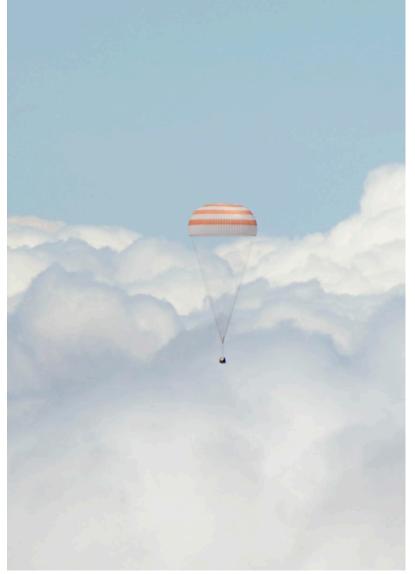
# Heat Load for Select Entry Vehicles

Ref. Davies, C., "Planetary Mission Entry Vehicles, Quick Reference Guide: Version 3"

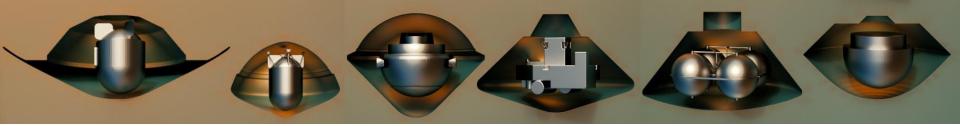


# Questions/Discussion?





NASA/JPL-Caltech/Univ. of Arizona ISS Expedition 31, NASA/Bill Ingalls International Planetary Probe Workshop 10 Short Course 2013



# Aerodynamics References

#### Introduction

Blunt Body Aerodynamics<sup>1</sup>

#### **Entry Vehicles**

CORONA Aerodynamics<sup>2–4</sup>

Early Manned Aerodynamics<sup>5</sup>

Apollo<sup>6–10</sup>

MPCV,Orion<sup>11–14</sup>

Viking Aerodynamics<sup>15–18</sup>

Mars Pathfinder<sup>19–21</sup>

Mars Exploration Rover<sup>22–24</sup>

Mars Science Laboroatory<sup>25–28</sup>

Mars Phoenix<sup>29</sup>

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Pioneer Venus<sup>31</sup>

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 $Genesis^{35} \\$ 

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Modified Newtonian<sup>40</sup>

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#### **Dynamic Stability**

Dynamic Stability Testing<sup>46–50</sup>
Dynamic Stability Theory and Computation<sup>51–56</sup>

#### **Reaction Control System**

Reaction Control Systems<sup>57–59</sup>

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